


Math 5421
An Introduction to
Mathematical Climate Models

Spring 2025
 1:25 – 3:20 Tuesdays and Thursdays
 Blegen Hall 155


Richard McGehee, Instructor
 458 Vincent Hall
 mcgehee@umn.edu
 www-users.cse.umn.edu/~mcgehee/

course website
<https://www-users.cse.umn.edu/~mcgehee/Course/Math5421/>




Math 5421 1/30/2025

2




Math 5421
Energy Balance

What determines Earth's surface temperature?



Math 5421 1/30/2025


3



Math 5421
Energy Balance


What determines Earth's surface temperature?

Conservation of Energy
 Heat is a form of energy.
 Temperature measures heat.



Math 5421 1/30/2025

4




Math 5421
Energy Balance

What determines Earth's surface temperature?

Conservation of Energy
 Heat is a form of energy.
 Temperature measures heat.


temperature change ~ energy in – energy out

↗ ↖
 short wave energy from the Sun long wave energy from the Earth



Math 5421 1/30/2025

5



Math 5421
Energy Balance


What determines Earth's surface temperature?

Conservation of Energy
 Heat is a form of energy.
 Temperature measures heat.

temperature change ~ energy in – energy out


↗ ↖
 short wave energy from the Sun long wave energy from the Earth

Everything else is detail.



Math 5421 1/30/2025

6

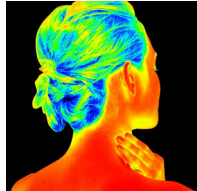


Math 5421
Energy Balance

Black-Body Radiation

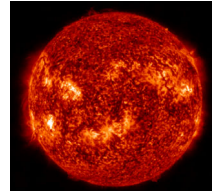
Every object emits electromagnetic radiation according to its temperature. The intensity of the radiation (power flux) is approximated by a theoretical object called a *perfect black body*.

human




<https://letstalkscience.ca/educational-resources/backgrounders/thermal-imaging>

sun



<https://solarsystem.nasa.gov/solar-system/sun/overview/>



Math 5421 1/30/2025

7

Math 5421
Energy Balance

Black-Body Radiation
Stefan-Boltzmann Law

watts per square meter
power flux (W/m²)

$F = \sigma T^4$

kelvin
temperature (K)
0 K = -273°C = "absolute zero"

Stefan-Boltzmann constant
 $\sigma \approx 5.67 \times 10^{-8} \text{ W/m}^2 \text{ K}^4$

watt = joule per second = "power"

8

Math 5421
Energy Balance

Black-Body Radiation
Stefan-Boltzmann Law

watts per square meter
power flux (W/m²)

$F = \sigma T^4$

kelvin
temperature (K)
0 K = -273°C = "absolute zero"

Stefan-Boltzmann constant
 $\sigma \approx 5.67 \times 10^{-8} \text{ W/m}^2 \text{ K}^4$

Reasonable approximation:
All bodies in the solar system radiate energy according to this law.

9

Math 5421
Energy Balance

Black-Body Radiation
Stefan-Boltzmann Law

watts per square meter
power flux (W/m²)

$F = \sigma T^4$

kelvin
temperature (K)
0 K = -273°C = "absolute zero"

Stefan-Boltzmann constant
 $\sigma \approx 5.67 \times 10^{-8} \text{ W/m}^2 \text{ K}^4$

Reasonable approximation:
All bodies in the solar system radiate energy according to this law.

Let's try the Sun.

10

Math 5421
Energy Balance

What is the surface temperature of the Sun?

11

Math 5421
Energy Balance

What is the surface temperature of the Sun?

5772K ≈ 10000°F

12

Math 5421
Stefan-Boltzmann Law

power flux (W/m²)

$F = \sigma T^4$

temperature (K)

Stefan-Boltzmann constant
 $\sigma \approx 5.67 \times 10^{-8} \text{ W/m}^2 \text{ K}^4$

surface temperature of the Sun: 5772K
power flux: $5.67 \times 10^{-8} \times (5772)^4 =$
 $6.29 \times 10^7 \text{ W/m}^2$

13


Math 5421
Stefan-Boltzman Law

$$F = \sigma T^4$$

power flux (W/m²) temperature (K)

Stefan-Boltzmann constant
 $\sigma \approx 5.67 \times 10^{-8} \text{ W/m}^2 \text{ K}^4$

surface temperature of the Sun: 5772K
power flux: $5.67 \times 10^{-8} \times (5772)^4 =$
Is this a lot? **$6.29 \times 10^7 \text{ W/m}^2$**



MCRN Math 5421 1/30/2025

14

Math 5421
Stefan-Boltzman Law

$$F = \sigma T^4$$


power flux (W/m²) temperature (K)

Stefan-Boltzmann constant
 $\sigma \approx 5.67 \times 10^{-8} \text{ W/m}^2 \text{ K}^4$

surface temperature of the Sun: 5772K
power flux: $5.67 \times 10^{-8} \times (5772)^4 =$
Is this a lot? **$6.29 \times 10^7 \text{ W/m}^2$**

Wikipedia: Prairie Island Nuclear Power Plant capacity =
522 + 519 = 1041 MW $\approx 10^9$ watts
 $10^9 / (6.29 \times 10^7) \approx 16 \text{ m}^2$
Sixteen square meters on the Sun's surface produces as much
power as a 1-gigawatt nuclear power plant.

https://en.wikipedia.org/wiki/Prairie_Island_Nuclear_Power_Plant



MCRN Math 5421 1/30/2025

15

Math 5421
Stefan-Boltzman Law


$$F = \sigma T^4$$

power flux (W/m²) temperature (K)

Stefan-Boltzmann constant
 $\sigma \approx 5.67 \times 10^{-8} \text{ W/m}^2 \text{ K}^4$

surface temperature of the Sun: 5772K
power flux: $5.67 \times 10^{-8} \times (5772)^4 =$
 $6.29 \times 10^7 \text{ W/m}^2$

What is the total energy output of the Sun?



MCRN Math 5421 1/30/2025

16

Math 5421
Stefan-Boltzman Law

$$F = \sigma T^4$$


power flux (W/m²) temperature (K)

Stefan-Boltzmann constant
 $\sigma \approx 5.67 \times 10^{-8} \text{ W/m}^2 \text{ K}^4$

surface temperature of the Sun: 5772K
power flux: $5.67 \times 10^{-8} \times (5772)^4 =$
 $6.29 \times 10^7 \text{ W/m}^2$

How many square meters on the surface of the Sun?

What is the total energy output of the Sun?



MCRN Math 5421 1/30/2025

17

Math 5421
Stefan-Boltzman Law

$$F = \sigma T^4$$

power flux (W/m²) temperature (K)


Stefan-Boltzmann constant
 $\sigma \approx 5.67 \times 10^{-8} \text{ W/m}^2 \text{ K}^4$

surface temperature of the Sun: 5772K
power flux: $5.67 \times 10^{-8} \times (5772)^4 =$
 $6.29 \times 10^7 \text{ W/m}^2$

How many square meters on the surface of the Sun?

What is the total energy output of the Sun?

radius of the sun = 6.96×10^8 meters



MCRN Math 5421 1/30/2025

18

Math 5421
Stefan-Boltzman Law

$$F = \sigma T^4$$


power flux (W/m²) temperature (K)

Stefan-Boltzmann constant
 $\sigma \approx 5.67 \times 10^{-8} \text{ W/m}^2 \text{ K}^4$

surface temperature of the Sun: 5772K
power flux: $5.67 \times 10^{-8} \times (5772)^4 =$
 $6.29 \times 10^7 \text{ W/m}^2$

total solar power output: $6.29 \times 10^7 \times 4\pi(r_s)^2$
where $r_s =$ radius of the sun = 6.96×10^8 m
total solar output: **$3.83 \times 10^{26} \text{ W}$**

Surface area of the Sun



MCRN Math 5421 1/30/2025

19

Math 5421
Stefan-Boltzman Law

$$F = \sigma T^4$$


power flux (W/m²) temperature (K)

Stefan-Boltzmann constant
 $\sigma \approx 5.67 \times 10^{-8} \text{ W/m}^2 \text{ K}^4$

surface temperature of the Sun: 5772K
power flux: $5.67 \times 10^{-8} \times (5772)^4 =$
 $6.29 \times 10^7 \text{ W/m}^2$

total solar power output: $6.29 \times 10^7 \times 4\pi(r_s)^2$,
where r_s = radius of the sun = $6.96 \times 10^8 \text{ m}$
total solar output: **$3.83 \times 10^{26} \text{ W}$**

How much is that?



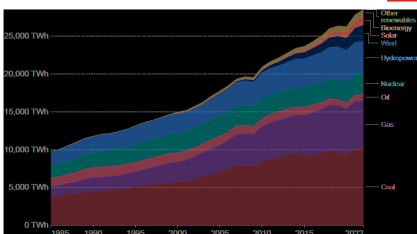
Math 5421 1/30/2025

20

Math 5421
Stefan-Boltzman Law

Global Annual Electricity Production
2022 production: about 28,000 TWh

Electricity production by source, World



28,000 TWh = $28 \times 10^{15} \text{ Wh}$
1 Wh = $3.6 \times 10^3 \text{ J}$
28,000 TWh = $28 \times 3.6 \times 10^{18} \text{ J} \approx 10^{20} \text{ J}$
Solar Output = **$3.83 \times 10^{26} \text{ J/s}$**
seconds of solar output = $10^{20} / 3.83 \times 10^{26} \approx$
 $0.26 \times 10^{-6} = 260 \times 10^{-9} =$
260 nanoseconds

Math 5421 1/30/2025

21

Math 5421
Stefan-Boltzman Law

$$F = \sigma T^4$$


power flux (W/m²) temperature (K)

Stefan-Boltzmann constant
 $\sigma \approx 5.67 \times 10^{-8} \text{ W/m}^2 \text{ K}^4$

surface temperature of the Sun: 5772K
power flux: $5.67 \times 10^{-8} \times (5772)^4 =$
 $6.29 \times 10^7 \text{ W/m}^2$

total solar power output: $6.29 \times 10^7 \times 4\pi(r_s)^2$,
where r_s = radius of the sun = $6.96 \times 10^8 \text{ m}$
total solar output: **$3.83 \times 10^{26} \text{ W}$**

260 nanoseconds = time it takes for the Sun to produce the equivalent of the annual global electricity production.



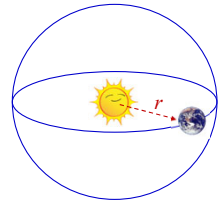
Math 5421 1/30/2025

22

Math 5421
Insolation

Incoming solar Radiation

How much energy from the Sun is hitting the Earth?



Math 5421 1/30/2025

23

Math 5421
Insolation

Incoming solar Radiation

How much energy from the Sun is hitting the Earth?

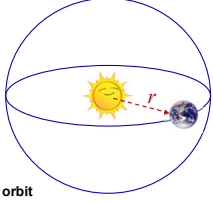
Solar flux at a distance r from the sun:

$$F = \frac{6.29 \times 10^7 \cdot 4\pi r_s^2}{4\pi r^2} = 6.29 \times 10^7 \left(\frac{r_s}{r}\right)^2 \text{ W/m}^2$$

$r_s = 7 \times 10^8 \text{ m}$
 $r = 1.5 \times 10^{11} \text{ m}$

$F = 1370 \text{ W/m}^2$

solar flux at Earth's orbit



<https://solarsystem.nasa.gov/planets/earth/overview/>

Math 5421 1/30/2025

24

Math 5421
Insolation

Incoming solar Radiation

How much energy from the Sun is hitting the Earth?

Solar flux at a distance r from the sun:

$$F = \frac{6.29 \times 10^7 \cdot 4\pi r_s^2}{4\pi r^2} = 6.29 \times 10^7 \left(\frac{r_s}{r}\right)^2 \text{ W/m}^2$$

$r_s = 7 \times 10^8 \text{ m}$
 $r = 1.5 \times 10^{11} \text{ m}$

$F = 1370 \text{ W/m}^2$

solar flux at Earth's orbit

Earth presents a disk to the Sun

area of disk: $F \times \pi r_e^2$ W

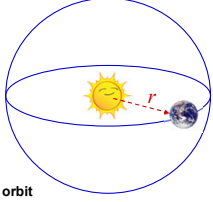
power intercepted by the Earth: $F \times \pi r_e^2$ W

Earth's surface area: $4\pi r_e^2$ m²

area of surface

Average surface flux: $\frac{F \times \pi r_e^2}{4\pi r_e^2} = \frac{F}{4} = 342 \text{ W/m}^2$

note the 4



Math 5421 1/30/2025

25

Math 5421
Insolation

Incoming solar Radiation

How much energy from the Sun is hitting the Earth?

Solar flux at a distance r from the sun:

$$F = \frac{6.29 \times 10^{27} 4\pi r_s^2}{4\pi r^2} = 6.29 \times 10^{27} \left(\frac{r_s}{r}\right)^2 \text{ W/m}^2$$

$r_s = 7 \times 10^8 \text{ m}$
 $r = 1.5 \times 10^{11} \text{ m}$

$F = 1370 \text{ W/m}^2$

solar flux at Earth's orbit

Note:
The IPCC uses $F = 1368$ and $F/4 = 342$, which are the values we will use.

Average surface flux: $\frac{F \times \pi r_E^2}{4\pi r_E^2} = \frac{F}{4} = 342 \text{ W/m}^2$

Math 5421 1/30/2025

26

Math 5421
Insolation

Incoming solar Radiation

How much energy from the Sun is hitting the Earth?

Solar flux at a distance r from the sun: $F = 1368 \text{ W/m}^2$

Power intercepted by Earth: $F \times \pi r_E^2 \text{ W}$

$r_E = \text{radius of Earth} = 6.37 \times 10^6 \text{ m}$

https://en.wikipedia.org/wiki/Earth_radius

Power intercepted by Earth:
 $F \times \pi r_E^2 = 1368 \times \pi \times 6.37^2 \times 10^{12} = 1.74 \times 10^{17} \text{ W}$

How much is that?

Math 5421 1/30/2025

27


Math 5421
Insolation

Incoming solar Radiation

How much energy from the Sun is hitting the Earth?

Solar power intercepted by Earth: $1.74 \times 10^{17} \text{ W}$

Biologically Stored Energy
total coal reserves: 10^{15} kg
energy content: $3 \times 10^7 \text{ J/kg}$
total energy in coal reserves: $3 \times 10^{22} \text{ J}$
How long: $3 \times 10^{22} / 1.74 \times 10^{17}$
 $\approx 1.7 \times 10^5 \text{ seconds}$
 $\approx 47 \text{ hours}$
 $\approx 2 \text{ days of insolation}$



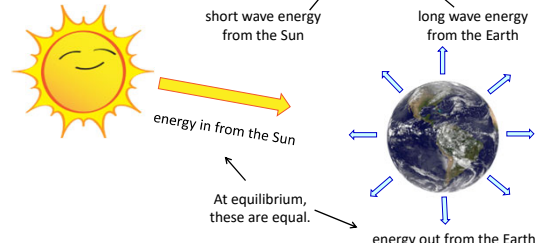
Math 5421 1/30/2025

28

Math 5421
Energy Balance

What determines the Earth's surface temperature?

temperature change \sim energy in - energy out



short wave energy from the Sun

long wave energy from the Earth

energy in from the Sun

energy out from the Earth

At equilibrium, these are equal.

Math 5421 1/30/2025

29

Math 5421
Energy Balance

What determines the Earth's surface temperature?

temperature change \sim energy in - energy out

energy in from the Sun

energy out from the Earth


Simple Model
Assume that Earth is a perfectly thermally conducting black body.

energy in from the Sun 342 W/m^2

energy out from the Earth $\sigma T^4 \text{ W/m}^2$

$$T = (342 / \sigma)^{1/4} = (342 / 5.67 \times 10^{-8})^{1/4} = 279\text{K} = 6^\circ\text{C} = 43^\circ\text{F}$$

Earth's global mean temperature: 57°F *Not bad!*



Math 5421 1/30/2025

30

Math 5421
Energy Balance

What determines the Earth's surface temperature?

Simple Model
Assume that Earth is a perfectly thermally conducting black body.

energy in from the Sun $Q = 342 \text{ W/m}^2$


energy out from the Earth $\sigma T^4 \text{ W/m}^2$

$$T = (342 / \sigma)^{1/4} = (342 / 5.67 \times 10^{-8})^{1/4} = 279\text{K} = 6^\circ\text{C} = 43^\circ\text{F}$$

Note that there is a differential equation lurking here.

temperature change \sim energy in - energy out

heat capacity $\rightarrow R \frac{dT}{dt} = Q - \sigma T^4$



Math 5421 1/30/2025

31

Math 5421
Energy Balance

temperature change ~ energy in - energy out

heat capacity $\rightarrow R \frac{dT}{dt} = Q - \sigma T^4$

Math 5421 1/30/2025

32

Math 5421
Goldilocks Zone

Goldilocks Zone

Earthling compatible life on exoplanets would be unlikely to occur unless there was the possibility of liquid water on the surface. This would mean the planet would not be too close to the star (too hot) or too far from the star (too cold). Instead, with an irresistible analogy with *Goldilocks and the Three Bears*, the temperature would be just right.

Math 5421 1/30/2025

33

Math 5421
Goldilocks Zone

Goldilocks Zone

Math 5421 1/30/2025

34

Math 5421
Goldilocks Zone

Goldilocks Zone

Math 5421 1/30/2025

35

Math 5421
Energy Balance

<https://www-users.cse.umn.edu/~mcgehee/Course/Math5421/assignments/A04.html>

Math 5421 1/30/2025

36