

An Introduction to Energy Balance Models

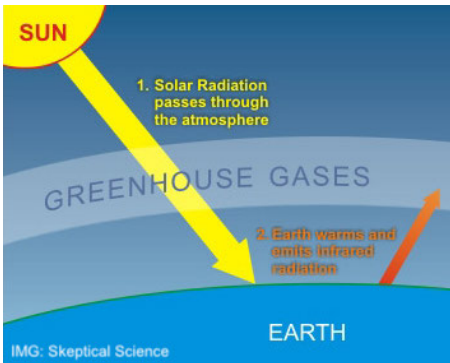
Alice Nadeau
(with a lot of slides from Dick McGehee)

University of Minnesota Mathematics of Climate Seminar

September 25, 2018

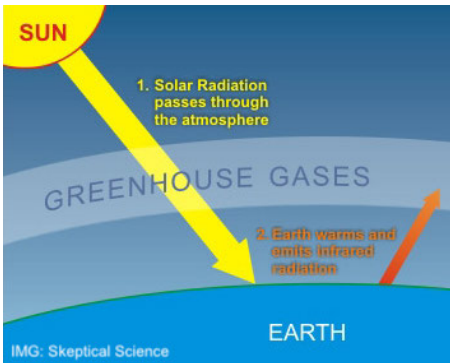
Conservation of energy

$$\text{temperature change} \sim \underbrace{\text{energy in}}_{\text{short wave radiation}} - \underbrace{\text{energy out}}_{\text{long wave radiation}}$$



Conservation of energy

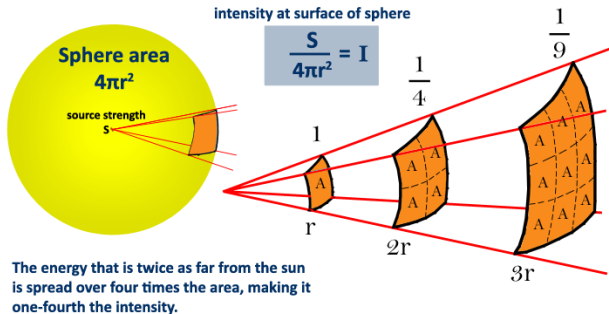
$$\text{temperature change} \sim \underbrace{\text{energy in}}_{\text{short wave radiation}} - \underbrace{\text{energy out}}_{\text{long wave radiation}}$$



Everything else is detail!

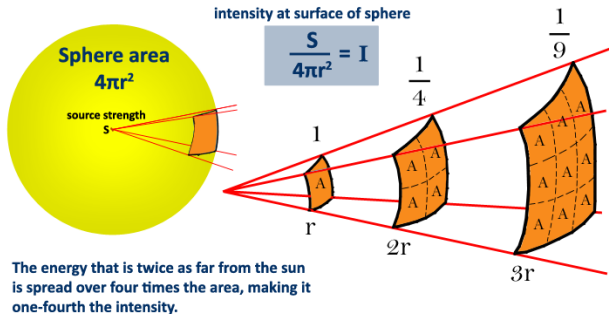
Initial Thoughts

Annual radiation from the Sun := Q

Finding Q 

IDEAS, USBC Geography Dept.

$$I_{\text{Earth}} = \frac{\text{power flux} \cdot \text{surface area}}{4\pi r_{\text{Earth}}^2} = \frac{(\sigma T_{\text{Sun}})^4 (4\pi r_{\text{Sun}}^2)}{4\pi r_{\text{Earth}}^2} \approx 1368 \text{ W m}^{-2}$$

Finding Q 

IDEAS, USBC Geography Dept.

$$I_{\text{Earth}} = \frac{\text{power flux} \cdot \text{surface area}}{4\pi r_{\text{Earth}}^2} = \frac{(\sigma T_{\text{Sun}})^4 (4\pi r_{\text{Sun}}^2)}{4\pi r_{\text{Earth}}^2} \approx 1368 \text{ W m}^{-2}$$

$$Q = \frac{I_{\text{Earth}} \cdot \pi r_{\text{Earth}}^2}{4\pi r_{\text{Earth}}^2} \approx 342 \text{ W m}^{-2}$$

Initial Thoughts

Annual radiation from the Sun := Q

Outgoing radiation := σT^4

→ Stefan-Boltzmann Law

Dynamical Models

Perfect thermally conducting black body:

$$R \frac{dT}{dt} = Q - \sigma T^4$$

Dynamical Models

Perfect thermally conducting black body:

$$R \frac{dT}{dt} = Q - \sigma T^4, \quad T^* = (Q/\sigma)^{1/4}$$

Dynamical Models

Perfect thermally conducting black body:

$$R \frac{dT}{dt} = Q - \sigma T^4, \quad T^* = (Q/\sigma)^{1/4}$$

Perfect thermally conducting black body **plus albedo**:

Albedo



Dynamical Models

Perfect thermally conducting black body:

$$R \frac{dT}{dt} = Q - \sigma T^4, \quad T^* = (Q/\sigma)^{1/4}$$

Perfect thermally conducting black body **plus albedo**:

$$R \frac{dT}{dt} = Q(1 - \alpha) - \sigma T^4, \quad T^* = ((1 - \alpha)Q/\sigma)^{1/4}$$

Dynamical Models for *Surface Temperature*

Convert to surface temperature:

$$R \frac{dT}{dt} = Q(1 - \alpha) - (A + BT), \quad T^* = ((1 - \alpha)Q - A)/B$$

Dynamical Models for *Surface Temperature*

Convert to surface temperature:

$$R \frac{dT}{dt} = Q(1 - \alpha) - (A + BT), \quad T^* = ((1 - \alpha)Q - A)/B$$

Include **latitude dependence**:

$$R \frac{\partial T}{\partial t} = Q_s(y)(1 - \alpha) - (A + BT(y, t)), \quad T^*(y) = ((1 - \alpha)Q_s(y) - A)/B$$

Dynamical Models for *Surface Temperature*

Convert to surface temperature:

$$R \frac{dT}{dt} = Q(1 - \alpha) - (A + BT), \quad T^* = ((1 - \alpha)Q - A)/B$$

Include **latitude dependence**:

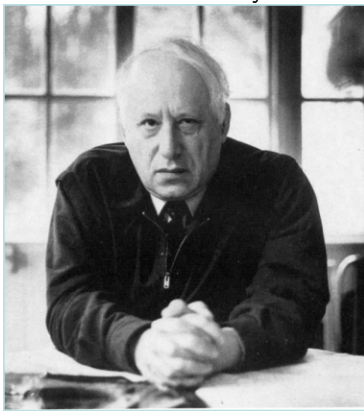
$$R \frac{\partial T}{\partial t} = Q_s(y)(1 - \alpha) - (A + BT(y, t)), \quad T^*(y) = ((1 - \alpha)Q_s(y) - A)/B$$

Include **heat transport**:

$$R \frac{\partial T}{\partial t} = Q_s(y)(1 - \alpha) - (A + BT(y, t)) - C \cdot f(T), \quad T^*(y) = \dots$$

Budyko vs. Sellers

Mikhail I. Budyko



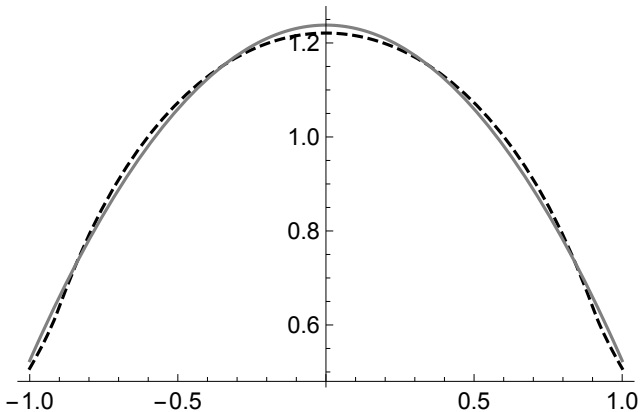
William D. Sellers



The Budyko Energy Balance Model

$$R \frac{\partial T}{\partial t} = \underbrace{Qs(y)(1 - \alpha)}_{\text{incoming radiation}} - \underbrace{(A + BT(y, t))}_{\text{OLR}} - \underbrace{C \left(T(y, t) - \overbrace{\bar{T}(t)}^{\text{ann. avg. temp.}} \right)}_{\text{heat transport}}$$

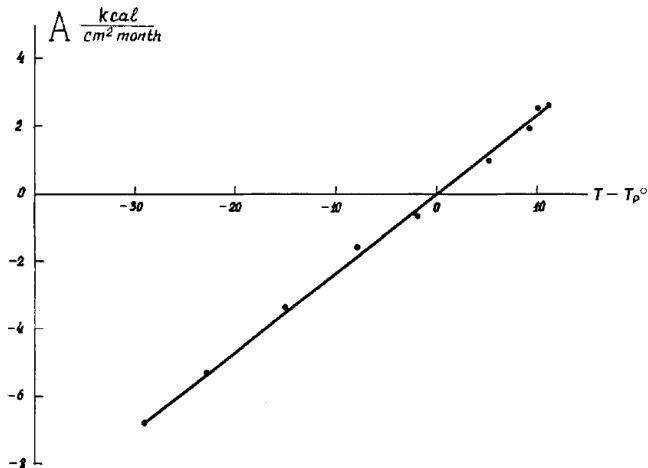
Incoming Solar Radiation Distribution: $s(y)$



Dashed: from first principles, Solid: Quadratic approximation

Finding A , B and C

A , B , and C are empirical parameters



M. Budyko, "The effects of solar radiation variations on the climate of the Earth," *Tellus*, 21: 1969, 611–619.

$$R \frac{\partial T}{\partial t} = \underbrace{Qs(y)}_{\text{insolation}} \underbrace{(1 - \alpha)}_{\text{albedo}} - \underbrace{(A + BT(y, t))}_{\text{OLR}} - \underbrace{C \left(T(y, t) - \overbrace{\bar{T}(t)}^{\int_0^1 T(y, t) dy} \right)}_{\text{heat transport}}$$

Incoming Solar Radiation Approximation:

$$s(y) \approx 1 - 0.238(3y^2 - 1)$$

Symmetry assumption:

$$\text{Equator} = 0 \leq y = \sin(\text{latitude}) \leq 1 = \text{North Pole}$$

Equilibrium Temperature Profile

$$0 = Q_s(y)(1 - \alpha) - (A + BT^*(y)) - C(T^*(y) - \bar{T}^*)$$

Equilibrium Temperature Profile

$$0 = Qs(y)(1 - \alpha) - (A + BT^*(y)) - C(T^*(y) - \bar{T}^*)$$

Integrate to find \bar{T}^* :

$$\begin{aligned} 0 &= \int_0^1 [Qs(y)(1 - \alpha) - (A + BT^*(y)) - C(T^*(y) - \bar{T}^*)] dy \\ &= Q \underbrace{\int_0^1 s(y) dy}_1 - Q \underbrace{\int_0^1 s(y) \alpha dy}_{\bar{\alpha}} - A \underbrace{\int_0^1 dy}_1 - B \underbrace{\int_0^1 T^*(y) dy}_{\bar{T}^*} \\ &\quad - C \underbrace{\int_0^1 T^*(y) dy + C \int_0^1 \bar{T}^* dy}_0 \\ &= Q(1 - \bar{\alpha}) - (A + B\bar{T}^*) \end{aligned}$$

⇒ **Equilibrium Global Mean Temperature:** $\bar{T}^* = \frac{1}{B}(Q(1 - \bar{\alpha}) - A)$

Equilibrium Temperature Profile

$$0 = Qs(y)(1 - \alpha) - (A + BT^*(y)) - C(T^*(y) - \overline{T^*})$$

$$\overline{T^*} = \frac{1}{B}(Q(1 - \bar{\alpha}) - A)$$

Equilibrium Temperature Profile

$$0 = Qs(y)(1 - \alpha) - (A + BT^*(y)) - C(T^*(y) - \bar{T}^*)$$

$$\bar{T}^* = \frac{1}{B} (Q(1 - \bar{\alpha}) - A)$$

Plug in \bar{T}^* and solve for $T^*(y)$:

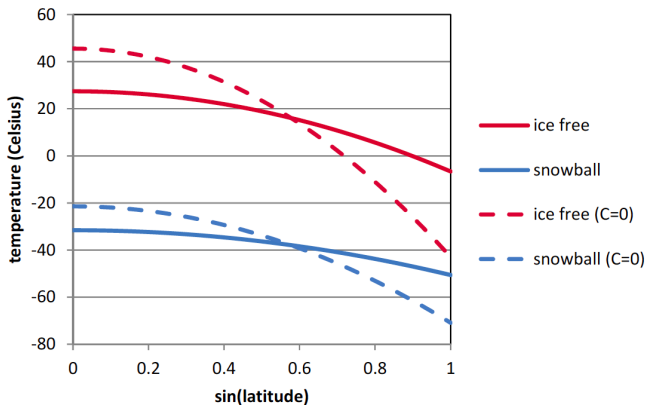
$$T^*(y) = \frac{1}{B + C} (Qs(y)(1 - \alpha) - A + C\bar{T}^*)$$

$$T^*(y) = \frac{1}{B + C} (Q_s(y)(1 - \alpha) - A + C\bar{T}^*)$$

$$\alpha = 0.32$$

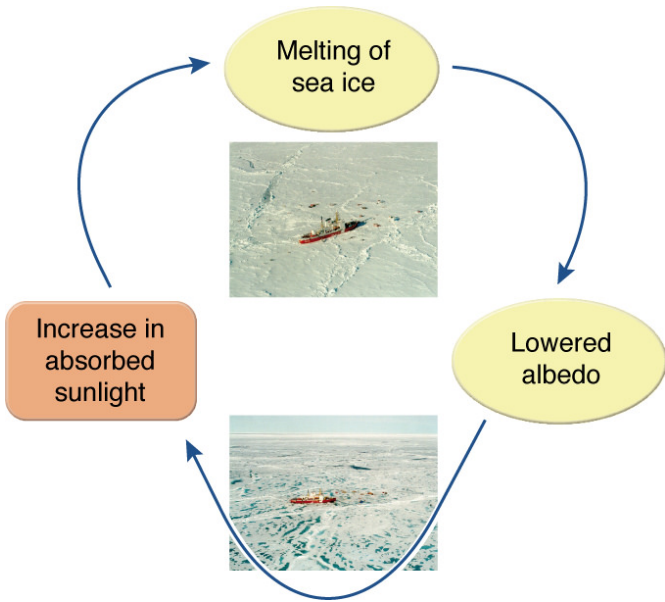
$$\alpha = 0.62$$

$$C = 3.04$$



From McGehee, Climate Seminar Sept. 19, 2017

Ice-Albedo Feedback



Non-uniform Albedo

$$R \frac{\partial T}{\partial t} = Q_s(y) (1 - \underbrace{\alpha(y, \eta)}) - (A + BT(y, t)) - C(T - \overline{T^*})$$

albedo depends on latitude

Non-uniform Albedo

$$R \frac{\partial T}{\partial t} = Q_s(y)(1 - \underbrace{\alpha(y, \eta)}) - (A + BT(y, t)) - C(T - \overline{T^*})$$

albedo depends on latitude

Ice Line Assumption: There is one ice line, η , in the northern hemisphere north of which there is always ice.

$$\alpha(y, \eta) = \begin{cases} \alpha_1 & 0 \leq y < \eta \\ \alpha_2 & \eta < y \leq 1 \end{cases}, \quad \alpha_1 < \alpha_2$$

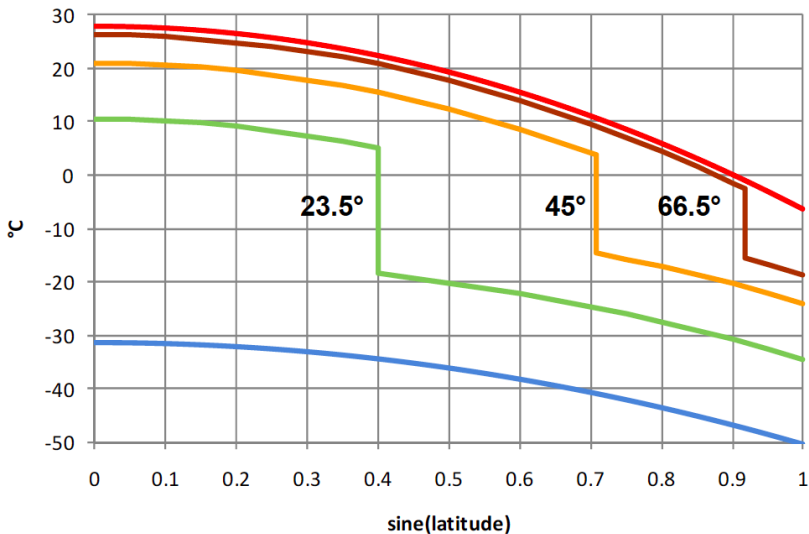
Equilibrium Temperature Profile depends on the Ice Line

$$T_{\eta}^*(y) = \frac{1}{B + C} (Qs(y)(1 - \alpha) - A + CT^*)$$
$$\bar{T}_{\eta}^* = \frac{1}{B} (Q(1 - \bar{\alpha}(\eta)) - A)$$

where

$$\bar{\alpha}(\eta) = \int_0^1 s(y)\alpha(y, \eta)dy = \alpha_1 \int_0^{\eta} s(y)dy + \alpha_2 \int_{\eta}^1 s(y)dy$$

Equilibrium Temperature Profile depends on the Ice Line



From McGehee, Climate Seminar Sept. 19, 2017

Dynamics of T

Experts only:

Theorem (Widiasih)

Let X be the space of functions where T lives and

$$L : X \rightarrow X; \quad LT := C\bar{T} - (B + C)T.$$

If $f(y) = Qs(y)(1 - \alpha(y, \eta)) - A$, then Budyko's equation can be written as a linear vector field on X :

$$R \frac{dT}{dt} = f + LT.$$

Furthermore, the operator L has only point spectrum, with all eigenvalues negative. Therefore all solutions are stable.

Dynamics of T

Experts only:

Theorem (Widiasih)

Let X be the space of functions where T lives and

$$L : X \rightarrow X; \quad LT := C\bar{T} - (B + C)T.$$

If $f(y) = Qs(y)(1 - \alpha(y, \eta)) - A$, then Budyko's equation can be written as a linear vector field on X :

$$R \frac{dT}{dt} = f + LT.$$

Furthermore, the operator L has only point spectrum, with all eigenvalues negative. Therefore all solutions are stable.

Everyone: For each fixed ice line η , there is a **globally stable** equilibrium solution for Budyko's equation.

Something seems wrong...

Something seems wrong...

Intuition:

- High temperature \Rightarrow ice melts \Rightarrow ice line moves north
- Low temperature \Rightarrow ice forms \Rightarrow ice line moves south

Something seems wrong...

Intuition:

- High temperature \Rightarrow ice melts \Rightarrow ice line moves north
- Low temperature \Rightarrow ice forms \Rightarrow ice line moves south

How do we model our intuitions?

Dynamic Ice Line

Ice Formation Assumption: Permanent ice forms if the annual average temperature is below $T_c = -10\text{ }^\circ\text{C}$ and melts if the annual average temperature is above T_c

Dynamic Ice Line

Ice Formation Assumption: Permanent ice forms if the annual average temperature is below $T_c = -10$ °C and melts if the annual average temperature is above T_c

$$\frac{d\eta}{dt} = \epsilon(T_\eta^*(\eta) - T_c)$$

Dynamics of the Ice Line

$$R \frac{\partial T}{\partial t} = Qs(y)(1 - \alpha(y, \eta)) - (A + BT(y, t)) - C(T - \overline{T}_\eta^*), \quad \frac{d\eta}{dt} = \epsilon(T_\eta^*(\eta) - T_c)$$

Experts only:

Theorem (Widiasih's Theorem)

For sufficiently small ϵ , the system has an attracting invariant curve given by the graph of a function $\Phi_\epsilon : [0, 1] \rightarrow X$. On this curve, the dynamics are approximated by the equation

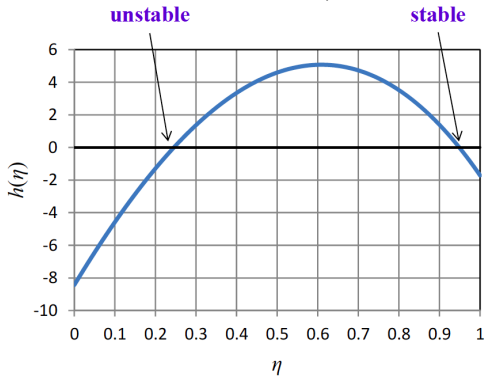
$$\frac{d\eta}{dt} = \epsilon(T_\eta^*(\eta) - T_c).$$

E. Widiasih, "Dynamics of the Budyko Energy Balance Model," *SIAM J. Appl. Dyn. Syst.*, 12(4), 2068–2092.

Dynamics of the Ice Line

$$R \frac{\partial T}{\partial t} = Qs(y)(1 - \alpha(y, \eta)) - (A + BT(y, t)) - C (T - \overline{T}_\eta^*), \quad \frac{d\eta}{dt} = \epsilon(T_\eta^*(\eta) - T_c)$$

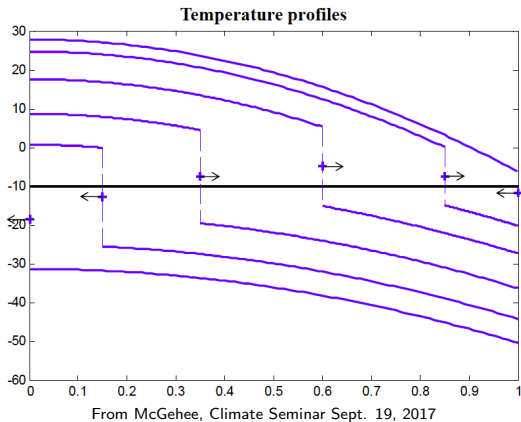
Everyone: $h(\eta) = T_\eta^*(\eta) - T_c$



From McGehee, Climate Seminar Sept. 19, 2017

The Budyko–Widiasih Model

$$R \frac{\partial T}{\partial t} = Qs(y)(1 - \alpha(y, \eta)) - (A + BT(y, t)) - C(T - \bar{T}_\eta^*), \quad \frac{d\eta}{dt} = \epsilon(T_\eta^*(\eta) - T_c)$$



Greenhouse Gasses in the Budyko–Widiasih Model

$$R \frac{\partial T}{\partial t} = Qs(y)(1 - \alpha(y, \eta)) - \underbrace{(A + BT(y, t))}_{\text{outgoing long wave radiation}} - C (T - \overline{T}_\eta^*)$$

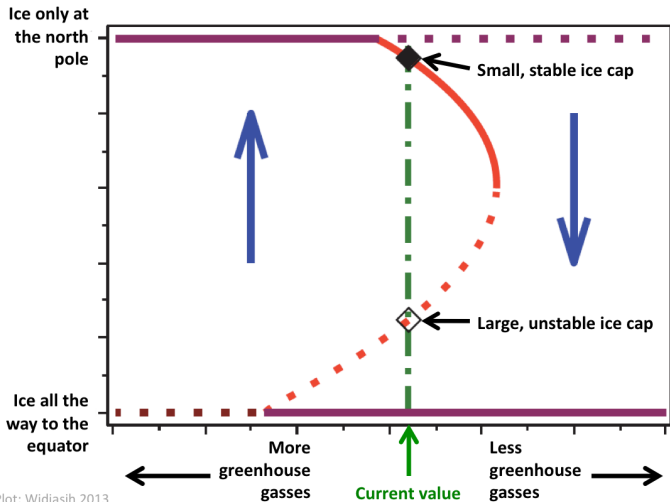
outgoing long wave radiation

$$\frac{d\eta}{dt} = \epsilon h(\eta, A)$$

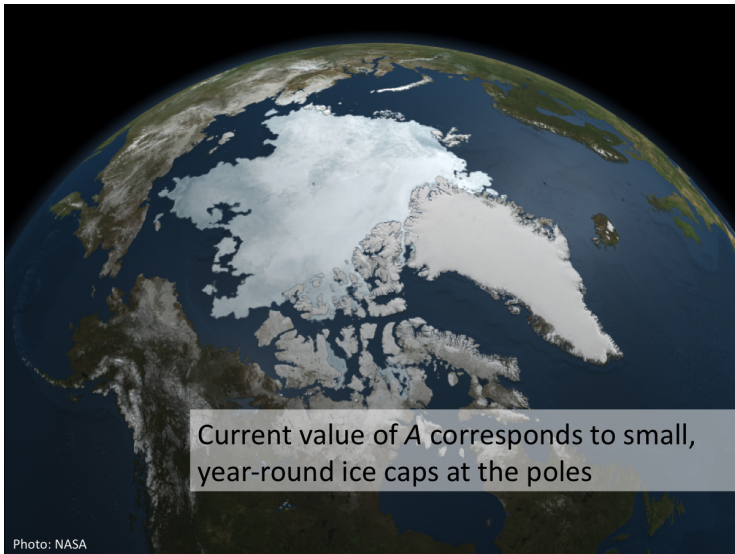
The parameter A is the **greenhouse gas parameter**.

Bifurcation Diagram for A

$$\frac{d\eta}{dt} = \epsilon h(\eta, A)$$



Current Earth

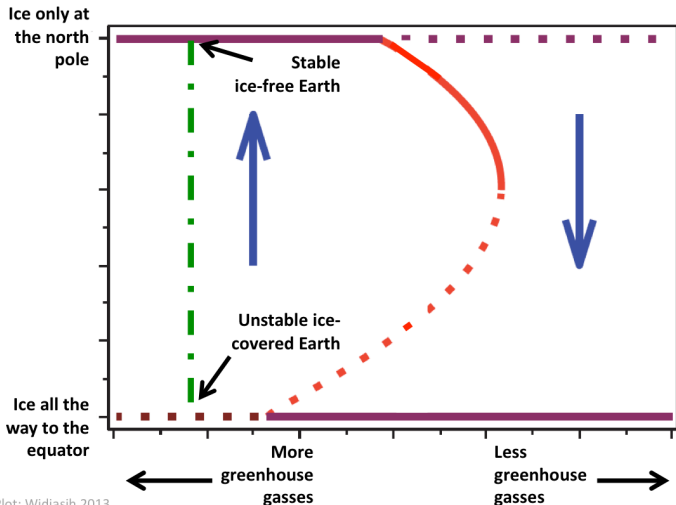


Current value of A corresponds to small, year-round ice caps at the poles

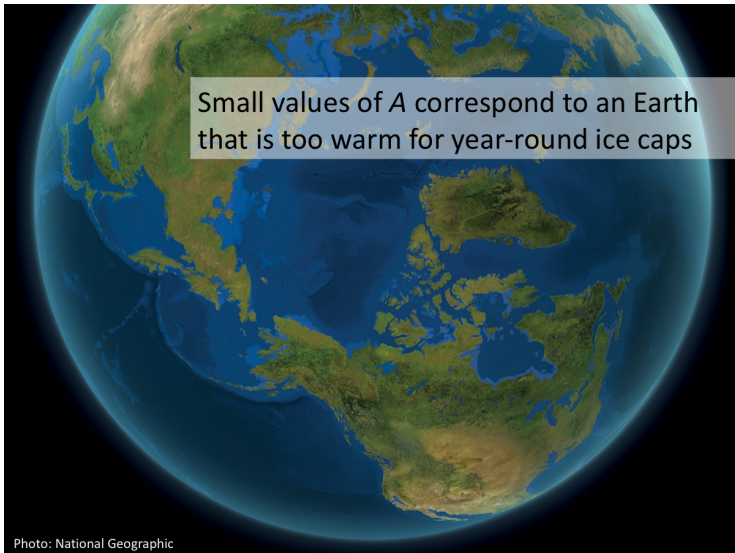
Photo: NASA

Bifurcation Diagram for A

$$\frac{d\eta}{dt} = \epsilon h(\eta, A)$$

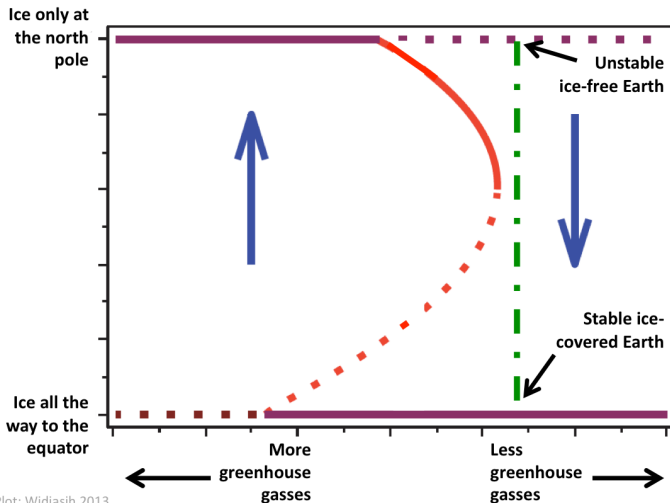


Future Earth?



Bifurcation Diagram for A

$$\frac{d\eta}{dt} = \epsilon h(\eta, A)$$



Past Earth?

Large values of A correspond to a completely ice covered Earth

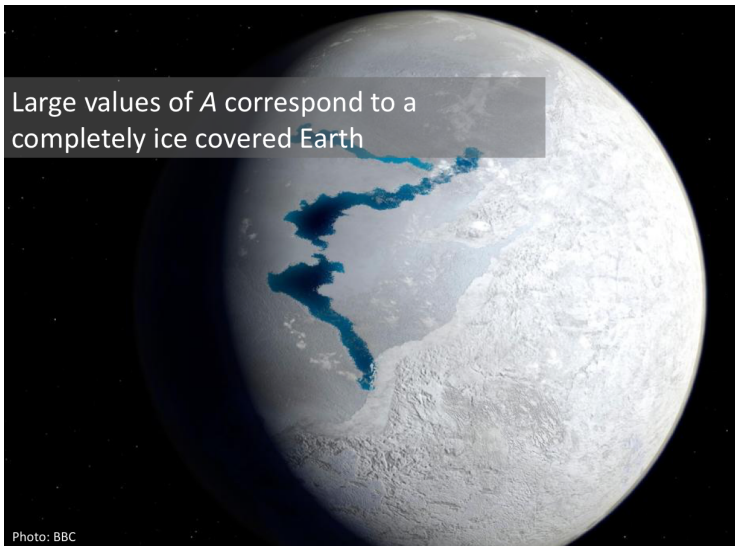


Photo: BBC

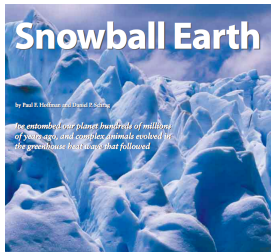
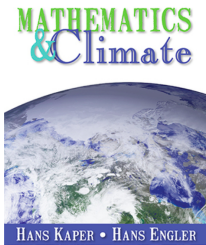
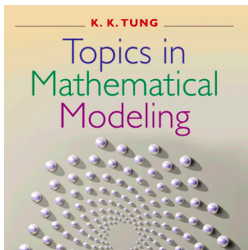
Evidence for Snowball Earth



Hoffman & Schrag, Snowball Earth, Scientific American, January 2000, 68-75

Further Reading

Everyone:



Experts:

- Barry, McGehee, Widiasih. (2017) "Nonsmooth Frameworks for and Extended Budyko Model."
- McGehee and Lehman. (2012) "A paleoclimate model of ice-albedo feedback forced by variations in Earth's orbit."
- McGehee and Widiasih. (2014) "A quadratic approximation to Budyko's ice-albedo feedback model with ice line dynamics."
- Walsh (2016) "Periodic orbits for a discontinuous vector field arising from a conceptual model of glacial cycles."
- Widiasih. (2013) "Dynamics of the Budyko Energy Balance Model."

Thank you!